

EXPERIMENTAL TEST OF PAULI EXCLUSION PRINCIPLE

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OUTLINE

- Introduction
 - Search for **anomalous carbon atoms**
 - Search for transitions to “**non-Paulian**” states using detector **NEMO-2**
 - Search for “**bosonic**” neutrino using **2β** decay data
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I. INTRODUCTION

- The Exclusion Principle was formulated by **W. Pauli in 1925**

(even before the concept of spin was introduced!)

- The **PEP** is valid for all identical particles with half-integral spin (fermions)
 - The **PEP** forbids that two identical fermions occupy the same quantum state
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THEORETICAL SITUATION

- ❑ **No satisfactory and consistent** mechanism of the **PEP** violation has been proposed so far
 - ❑ **There is no** consistent theoretical framework within which the sensitivities of different experiments can be compared
 - ❑ The exceptional place occupied by the **PEP** in modern physics does not imply that it does not need further painstaking experimental tests
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EXPERIMENT

- Search for **non-Paulian** atoms and nuclei
 - Search for transitions to **non-Paulian states** in atoms and nuclei
 - Check the statistics of **neutrinos**
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1. Search for anomalous atoms

- “**Non-Paulian**” atoms could be a cosmological origin, if not all 10^{80} electrons in the universe are antisymmetrized or if spontaneous transitions of ordinary atoms into “**non-Paulian**” atoms are possible.
 - The chemical properties of atoms with three electrons per **1s** shell must be similar to the properties of their “lower-order” neighbors in the periodic table (*for example, “non-Paulian” carbon is similar to boron*).
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In 1989 Novikov and Pomansky proposed to search for anomalous atoms which arose in the universe during the nucleosynthesis. B-C and F-Ne were proposed as most promising candidates.

1) During the formation of the solar system, a differentiation of the original homogeneous protomatter occurred, according to the chemical properties of the elements. And finally, each element (Z) contains anomalous atoms of the element (Z+1). For example, boron will contain anomalous atoms of carbon.

$$\hat{C}/B = \hat{C}/C \times P(Z+1)/P(Z)$$

$$P(Z+1)/P(Z) = 2.18 \cdot 10^6 \text{ for B-C}$$
$$= 3300 \text{ for F-Ne}$$

$$2) C = t \cdot P(Z+1) / \tau \cdot P(Z)$$

C is concentration of anomalous atoms in the host substance;

t is $\sim 4.5 \cdot 10^9$ y (average cosmological time interval during which anomalous atoms were in protomatter);

τ is the lifetime of an atomic electron with respect to violation of the PEP;

P(Z+1) and P(Z) are the cosmic abundance of elements.

Search for anomalous carbon atoms (JETP Lett. 68 (1998) 112)

□ $P(C)/P(B) = 2.18 \cdot 10^6$

□ B: ^{10}B (19%) and ^{11}B (81%)

□ C: ^{12}C (99%) and ^{13}C (1%)

The anomalous atom $^{12}\hat{\text{C}}$ contains **three K-shell** electrons and behaves chemically like a boron atom

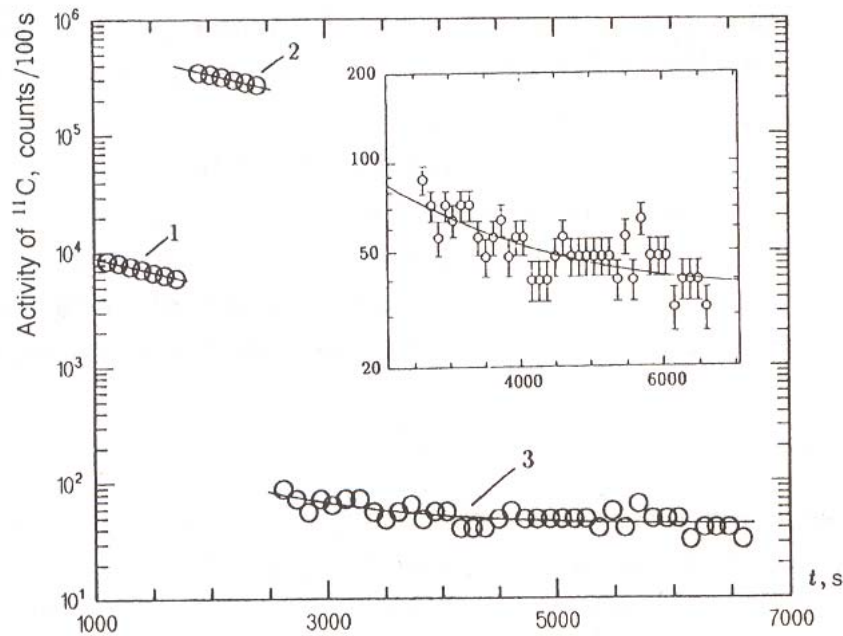
The idea of the experiment is to remove carbon atoms chemically from a boron sample and then to measure the content of carbon nuclei in it.

- The search for $\hat{\text{C}}$ was conducted by γ -**activation** analysis (this method does not produce any radioactive isotopes in B)
 $^{12}\text{C}(\gamma, n)^{11}\text{C}; \quad ^{11}\text{C}(20.38 \text{ m}) \rightarrow ^{11}\text{B} + \beta^+ + \nu$
 - The experiment was performed on the microtron at the **Institute of Physics Problems of the Russian Academy of Science**
 - Boron grown by zone melting was used (initial content of carbon is $1.5 \cdot 10^{-3} \text{ g/g}$)
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Procedure

- A boron sample (~ 0.1 g) was irradiated with γ -ray bramsstrahlung from 28 MeV electrons
 - After irradiation the boron sample was purified chemically from carbon
 - Activity of ^{11}C nuclei remaining in the boric acid solution was measured as a function of time (using γ - γ coincidence method)
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^{11}C activity versus time



1 - carbon separated from 0.1 g of boron after irradiation

2 - 0.006 g of carbon after irradiation (for comparison)

3 - boron (after chemical separation of carbon)

Inset: the detected positron activity in the boron solution.

There is a “positive” effect!!!

(Of course, it can be connected with rest of normal carbon in the boron sample.)

RESULTS

- Concentration of anomalous carbon atoms in boron is $\leq 5 \cdot 10^{-6} \text{ g/g}$



- Concentration of anomalous carbon atoms in carbon is $^{12}\hat{\text{C}}/^{12}\text{C} \leq 2.5 \cdot 10^{-12}$

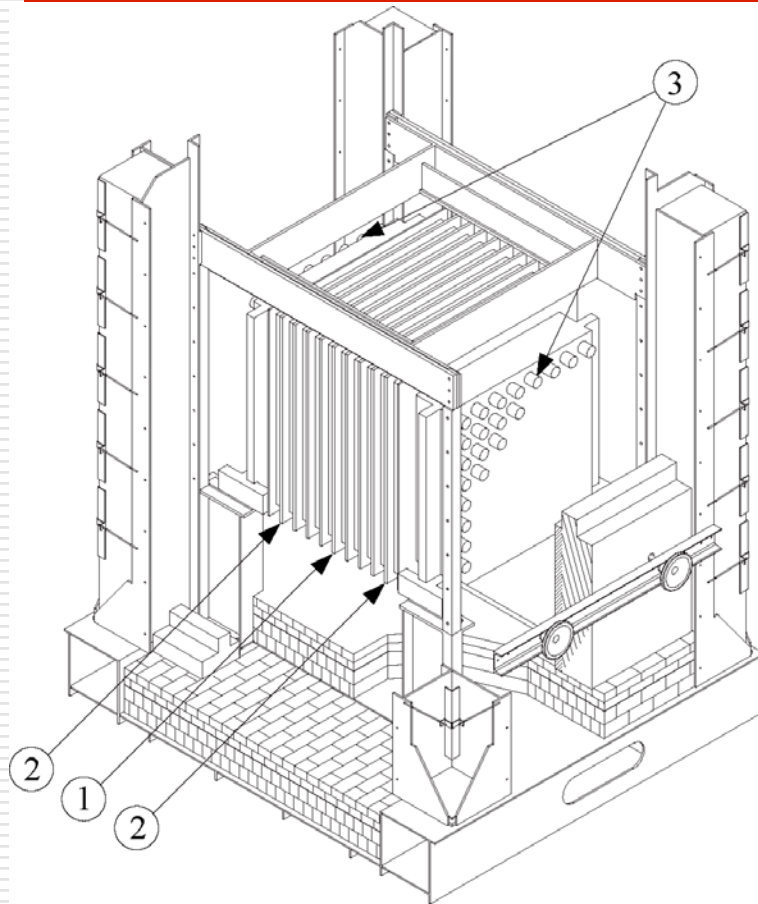


- The lifetime of electrons in a carbon atom relative to violation of the PEP $\tau \geq 2 \cdot 10^{21} \text{ y}$
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Best present limits

Be'/Be	$< 9 \cdot 10^{-12}$	Phys. Rev. Lett 85 (2000) 2701
$^{12}\hat{\text{C}}/^{12}\text{C}$	$\leq 2.5 \cdot 10^{-12}$	JETP Lett. 68 (1998) 112
$^{20}\tilde{\text{Ne}}/^{20}\text{Ne}$	$< 2 \cdot 10^{-21}$	Phys. Lett. B 240 (1990) 227
$^{36}\tilde{\text{Ar}}/^{36}\text{Ar}$	$< 4 \cdot 10^{-17}$	Phys. Lett. B 240 (1990) 227

2. Search for “non-Paulian” transitions using detector NEMO-2



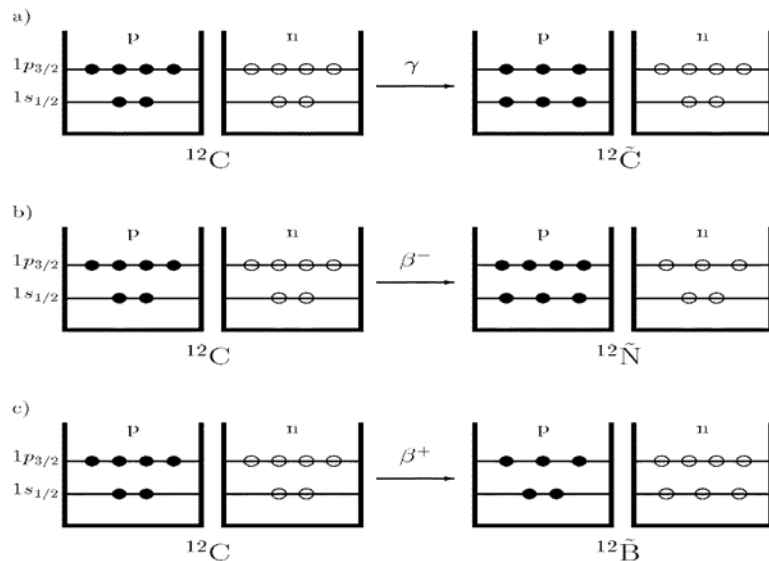
1 – Central frame with the source plane

2 – tracking device

3 – two scintillator walls

Full mass of ^{12}C is 170 kg

Non-Paulian transitions



Non-Paulian process with high energy γ -quantum emission ($E_\gamma \approx 18$ MeV)

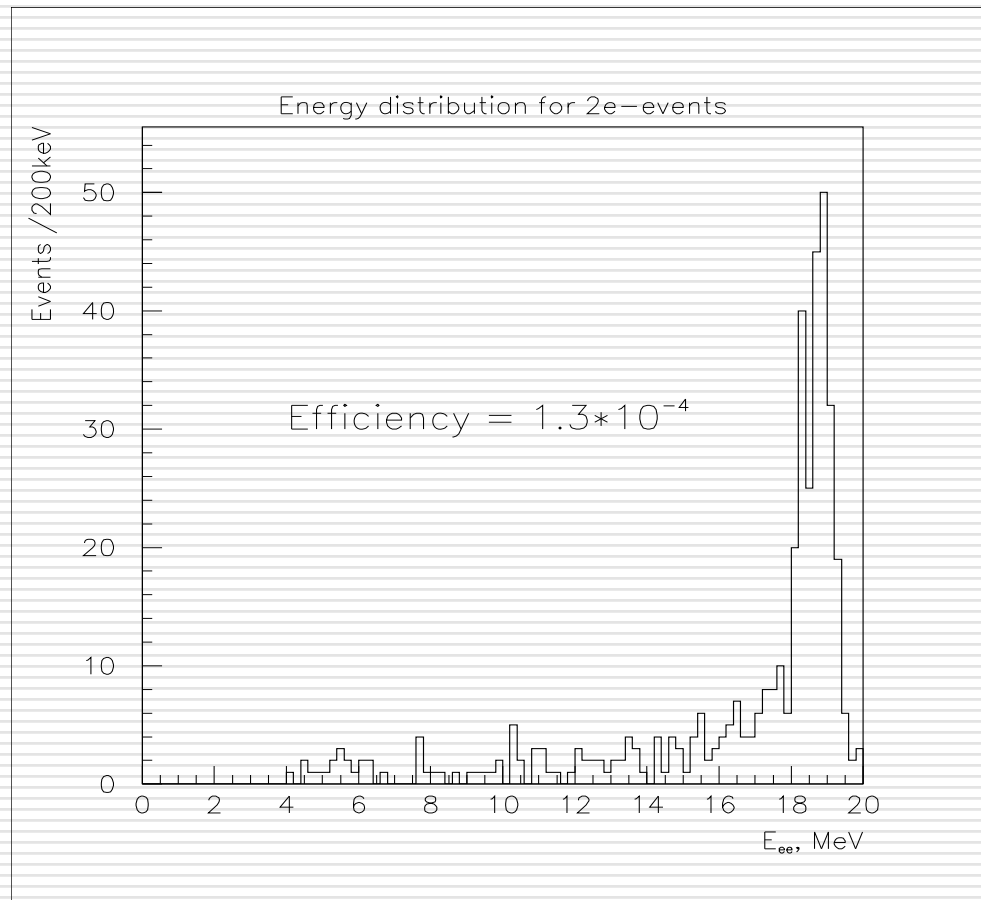
Non-Paulian β^- decay

($Q_\beta \approx 19$ MeV)

Non-Paulian β^+ decay

($Q_\beta \approx 17$ MeV)

Non-Paulian process with high energy γ -quantum emission



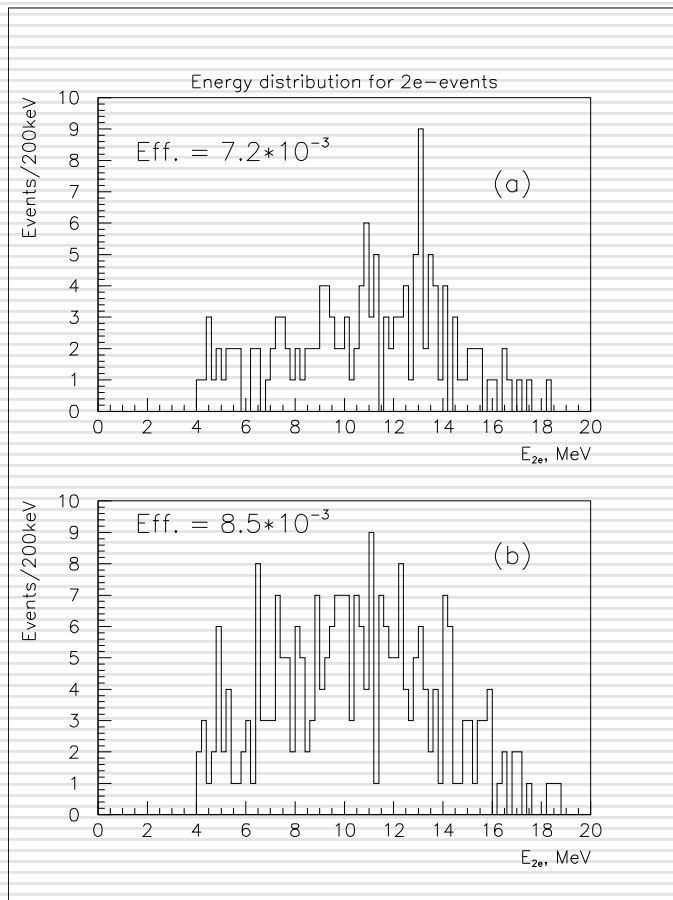
We are looking for e^+e^- pair created in the source.

Efficiency was calculated by MC.

No events with two tracks and summed energy **>4 MeV** were found after **16945 h** of measurements

$$T_{1/2} > 4.2 \cdot 10^{24} \text{ y at 90\% C.L.}$$

β^\pm decays to non-Paulian states



β^+

β^-

We are looking for **two tracks events**.

Selection: electron appears in a plastic scintillator, cross the tracking volume and source and enter a plastic scintillator on the opposite side of the detector.

Efficiency was calculated by MC

One event with energy > 4 **MeV** was found (2163 h with antineutron shield)

$T_{1/2}(\beta^+) > 2.6 \cdot 10^{24}$ y and $T_{1/2}(\beta^-) > 3.1 \cdot 10^{24}$ y at 90% C.L.

RESULTS (Eur. Phys. J. A 6 (1999) 361)

Channel	γ emission	β^-	β^+
Energy window (MeV)	[4,20]	[4,20]	[4,20]
Number of events	0	1	1
Efficiency	$1.3 \cdot 10^{-4}$	$8.5 \cdot 10^{-3}$	$7.2 \cdot 10^{-3}$
$T_{1/2}$, y (90% CL)	$> 4.2 \cdot 10^{24}$	$> 3.1 \cdot 10^{24}$	$> 2.6 \cdot 10^{24}$

Best previous $> 1.3 \cdot 10^{20}$ [1] $> 8 \cdot 10^{27}$ [2] $> 8 \cdot 10^{27}$ [2]

results

[1] B.A. Logan and A. Ljubicic, PR C20 (1979) 1957.

[2] D. Kekez, A. Ljubicic, B.A. Logan, Nature 348 (1990) 224.

Borexino CTF in 2004: $> 2.1 \cdot 10^{27} \text{ y}$
($^{12}\text{C} \rightarrow ^{12}\hat{\text{C}} + \gamma$)

□ Sensitivity of **NEMO-3**:
 $\sim 10^{27} \text{ y}$

□ Sensitivity of **SuperNEMO**:
 $\sim 10^{28} \text{ y}$

3. Search for “bosonic” neutrino using 2β decay data

- In 2005 **Dolgov and Smirnov** assumed that the **PEP** is violated for neutrinos and, consequently, neutrinos obey (at least partly) the Bose-Einstein statistics [PL B 621 (2005) 1]
 - **Consequences of this assumption:**
 - a) neutrinos may form cosmological Bose condensate (**dark matter!**)
 - b) “wrong” statistic of neutrinos could modify Big Bang nucleosynthesis
 - c) spectra of the supernova neutrinos may be changed
 - d) **PEP** violation for neutrinos can be tested in the two neutrino **double beta decay** experiments
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Why it can be possible for neutrinos?

- ❑ PEP never was checked for neutrinos
 - ❑ Neutrinos are only known neutral leptons
 - ❑ Neutrino can be a Majorana particle ($\nu \equiv \nu'$) and violate lepton number conservation
 - ❑ Neutrino has a very small mass
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If neutrino is bosonic (or partially bosonic) particle one can see the effect in $2\beta(2\nu)$ decay:

- Probability of decay will be changed
- Sum and single electron energy spectra will be changed
- Angular distribution will be changed

So, $2\beta(2\nu)$ decay is nice process to check possible PEP violation in neutrino sector

**A.S. Barabash, A.D. Dolgov, R. Dvornicky, F. Simkovic,
A.Yu. Smirnov, Nucl. Phys. B 783 (2007) 90.**

- The amplitude of the $2\beta(2\nu)$ decay can be parametrized as

$$A_{2\beta} = \cos^2\chi A_f + \sin^2\chi A_b$$

- The probability of $2\beta(2\nu)$ decay is equal to:

$$W_{\text{tot}} = \cos^4\chi W_f + \sin^4\chi W_b$$

where $W_{f,b} \sim |A_{f,b}|^2$

Predictions for decay rate:

□ ^{100}Mo :

Theory: $T_{1/2}^f(0^+_{\text{g.s.}}) = (6.8 \pm 3.4) \cdot 10^{18} \text{ y}$, $T_{1/2}^b(0^+_{\text{g.s.}}) = 8.9 \cdot 10^{19} \text{ y}$

Experiment: $T_{1/2}(0^+_{\text{g.s.}}) = (7.1 \pm 0.4) \cdot 10^{18} \text{ y}$,

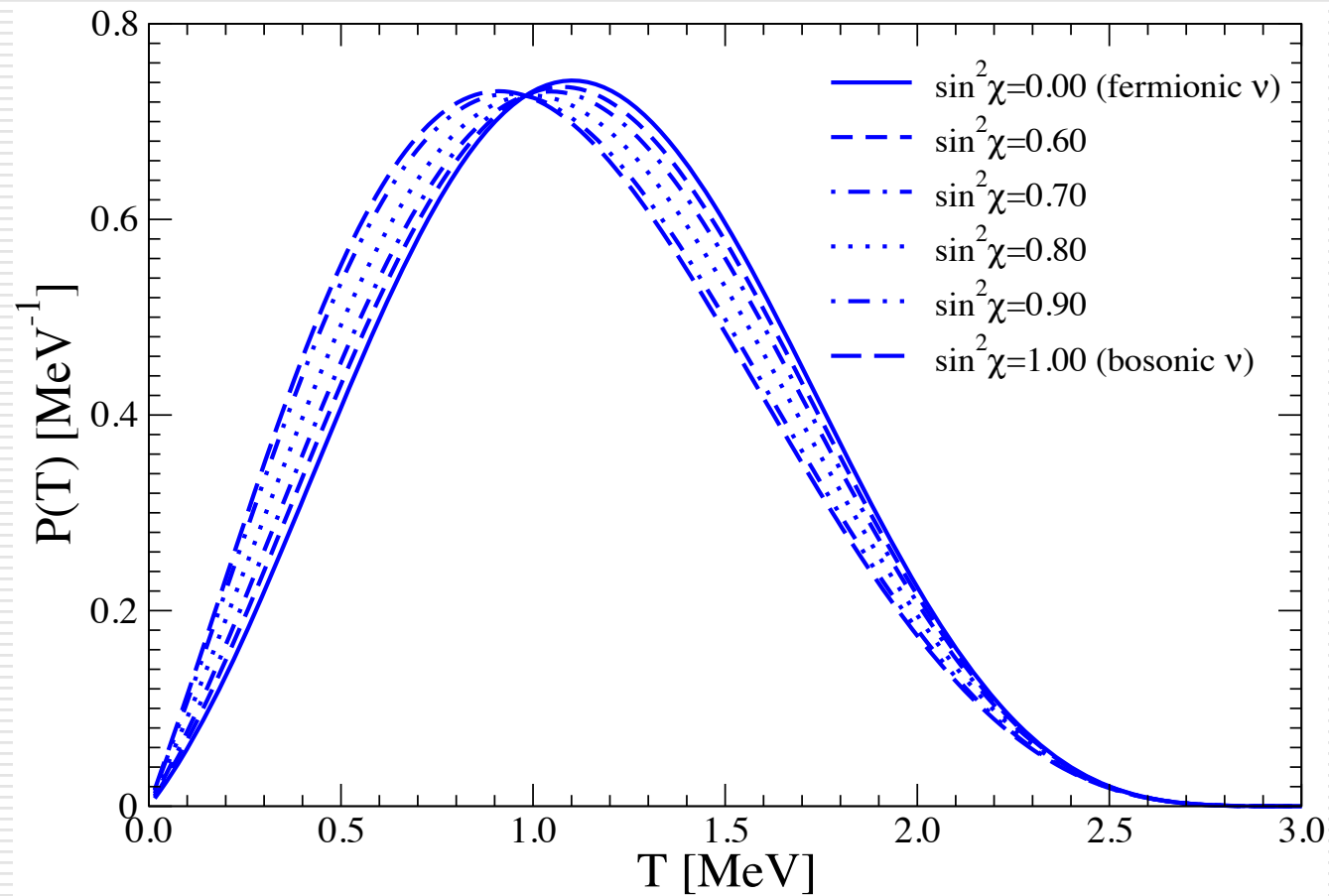
□ ^{76}Ge :

Theory: $T_{1/2}^f(0^+_{\text{g.s.}}) = (0.8 - 1.4) \cdot 10^{21} \text{ y}$, $T_{1/2}^b(0^+_{\text{g.s.}}) = 1.5 \cdot 10^{24} \text{ y}$

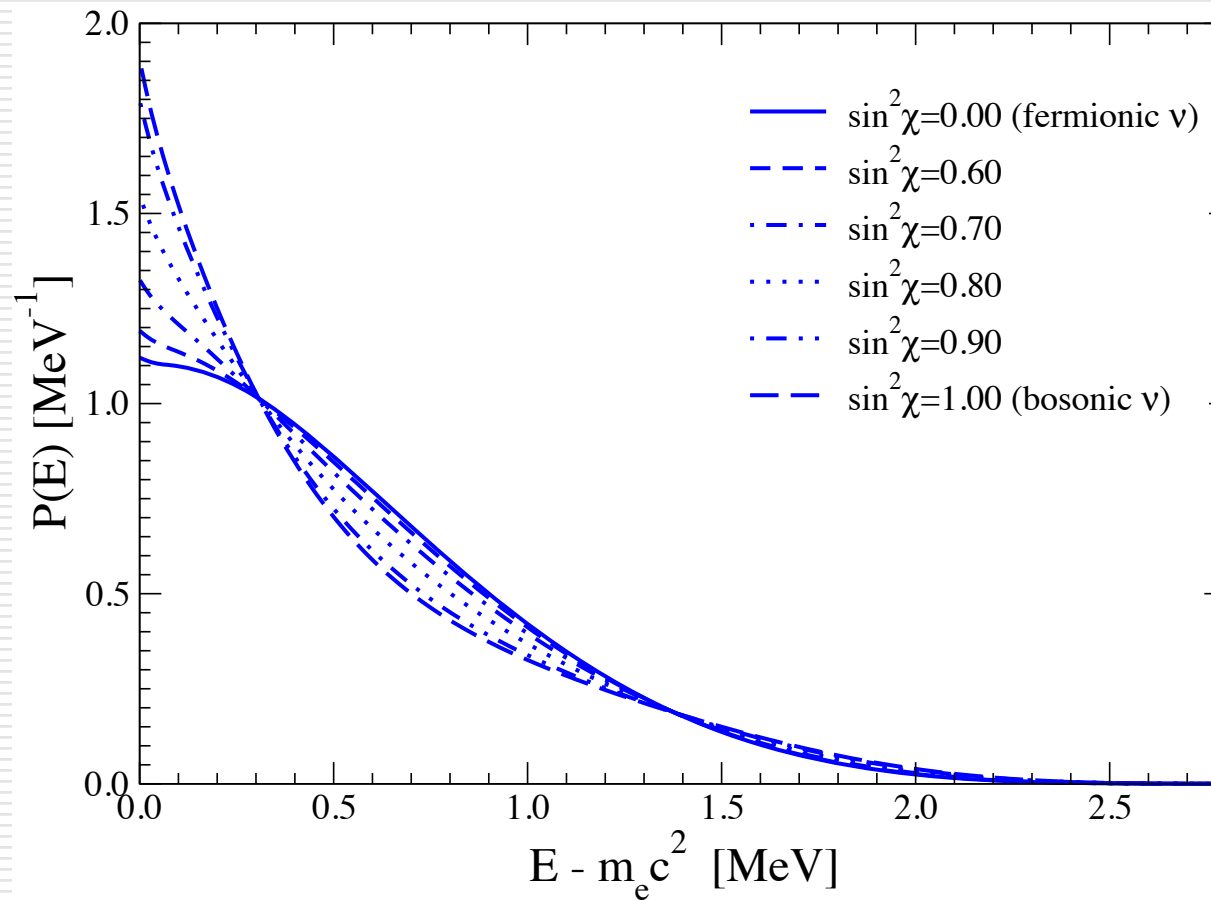
Experiment: $T_{1/2}(0^+_{\text{g.s.}}) = (1.5 \pm 0.1) \cdot 10^{21} \text{ y}$,

Conclusion: $\sin^2\chi < 0.5$

The normalized distribution of the total energy of two electrons

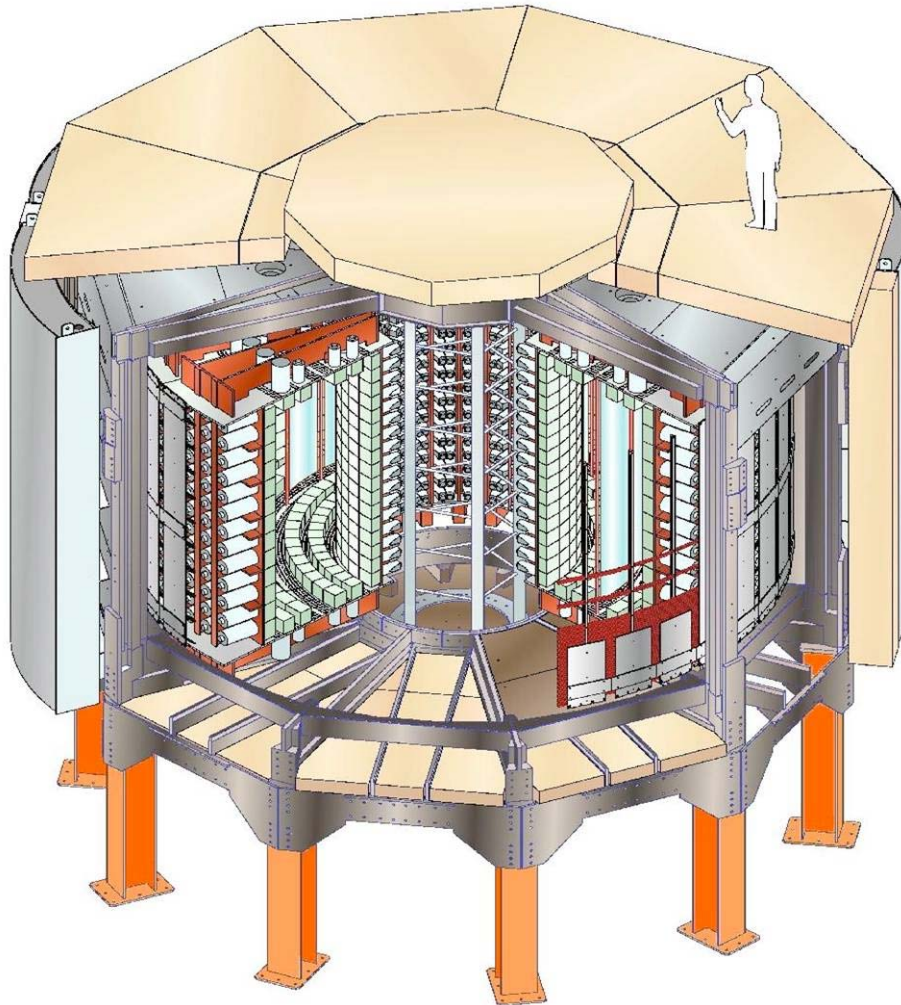


The normalized distribution of the single electron energy



The NEMO3 detector

Fréjus Underground Laboratory : 4800 m.w.e.



Source: 10 kg of $\beta\beta$ isotopes
cylindrical, $S = 20 \text{ m}^2$, 60 mg/cm^2

Tracking detector:

drift wire chamber operating
in Geiger mode (6180 cells)

Gas: He + 4% ethyl alcohol + 1% Ar + 0.1% H₂O

Calorimeter:

1940 plastic scintillators
coupled to low radioactivity PMTs

Magnetic field: 25 Gauss

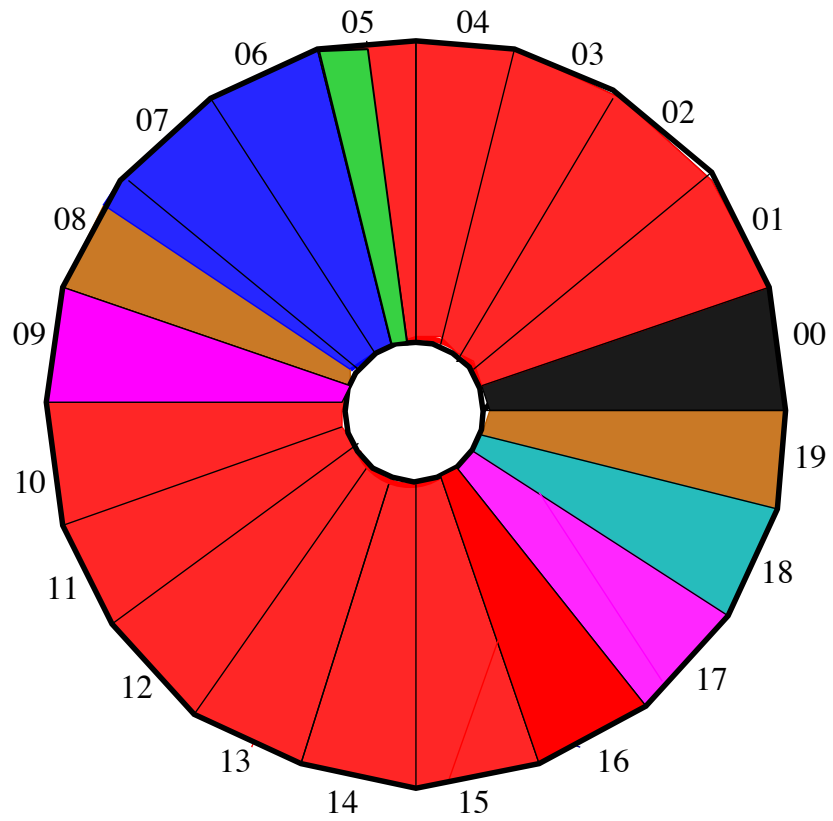
Gamma shield: Pure Iron (18 cm)

Neutron shield: borated water (~30 cm) + Wood (Top/Bottom/Gaps between water tanks)



Able to identify e^- , e^+ , γ and α

ββ decay isotopes in NEMO-3 detector



ββ2ν measurement

- ↑ **¹¹⁶Cd** **405 g**
 $Q_{\beta\beta} = 2805 \text{ keV}$
- ⁹⁶Zr** **9.4 g**
 $Q_{\beta\beta} = 3350 \text{ keV}$
- ¹⁵⁰Nd** **37.0 g**
 $Q_{\beta\beta} = 3367 \text{ keV}$
- ⁴⁸Ca** **7.0 g**
 $Q_{\beta\beta} = 4272 \text{ keV}$
- ¹³⁰Te** **454 g**
 $Q_{\beta\beta} = 2529 \text{ keV}$
- natTe** **491 g**
- Cu** **621 g**

External bkg measurement

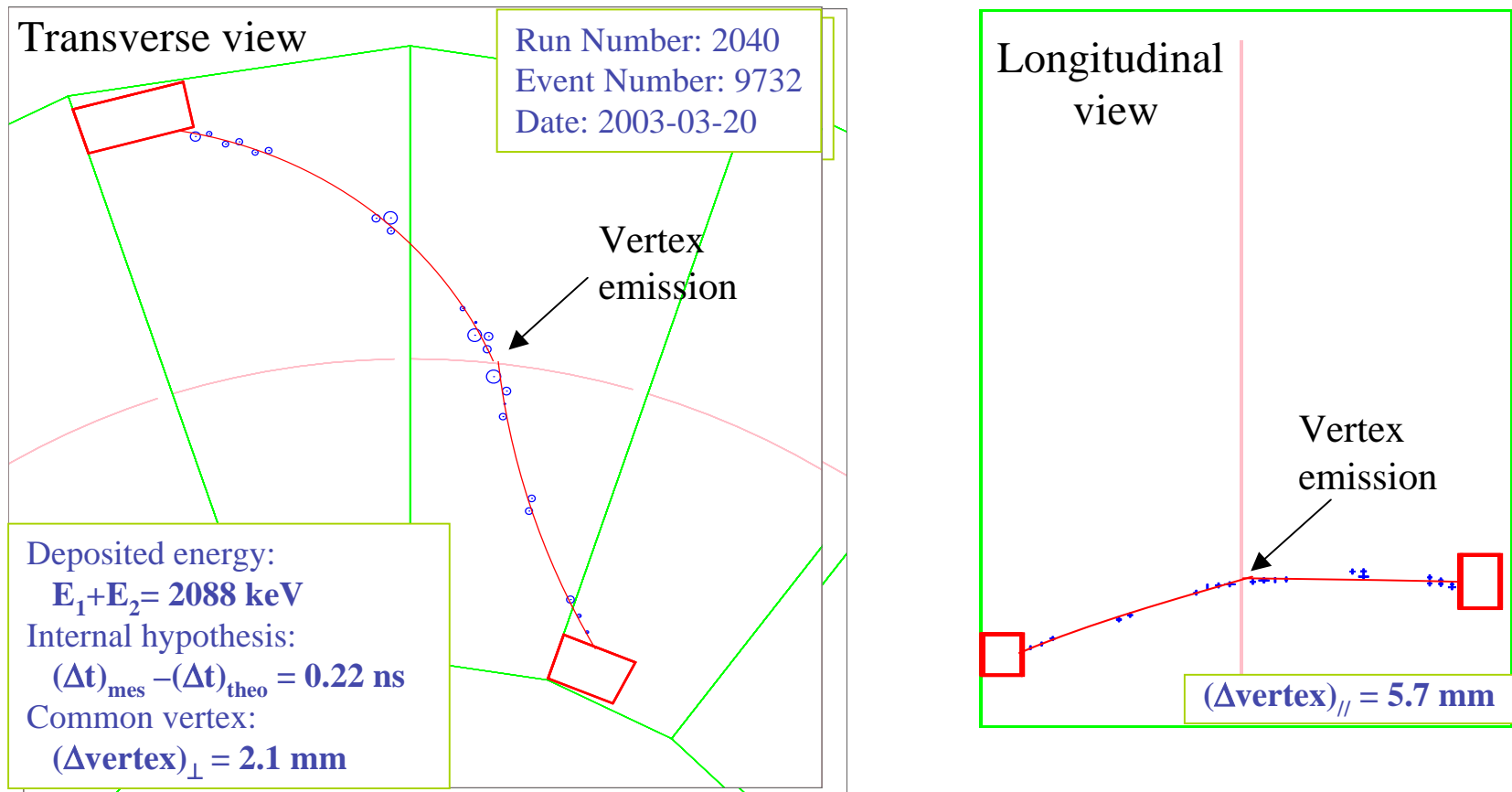
¹⁰⁰Mo **6.914 kg** **⁸²Se** **0.932 kg**
 $Q_{\beta\beta} = 3034 \text{ keV}$ $Q_{\beta\beta} = 2995 \text{ keV}$

ββ0ν search

(All enriched isotopes produced in Russia)

$\beta\beta$ events selection in NEMO-3

Typical $\beta\beta 2\nu$ event observed from ^{100}Mo

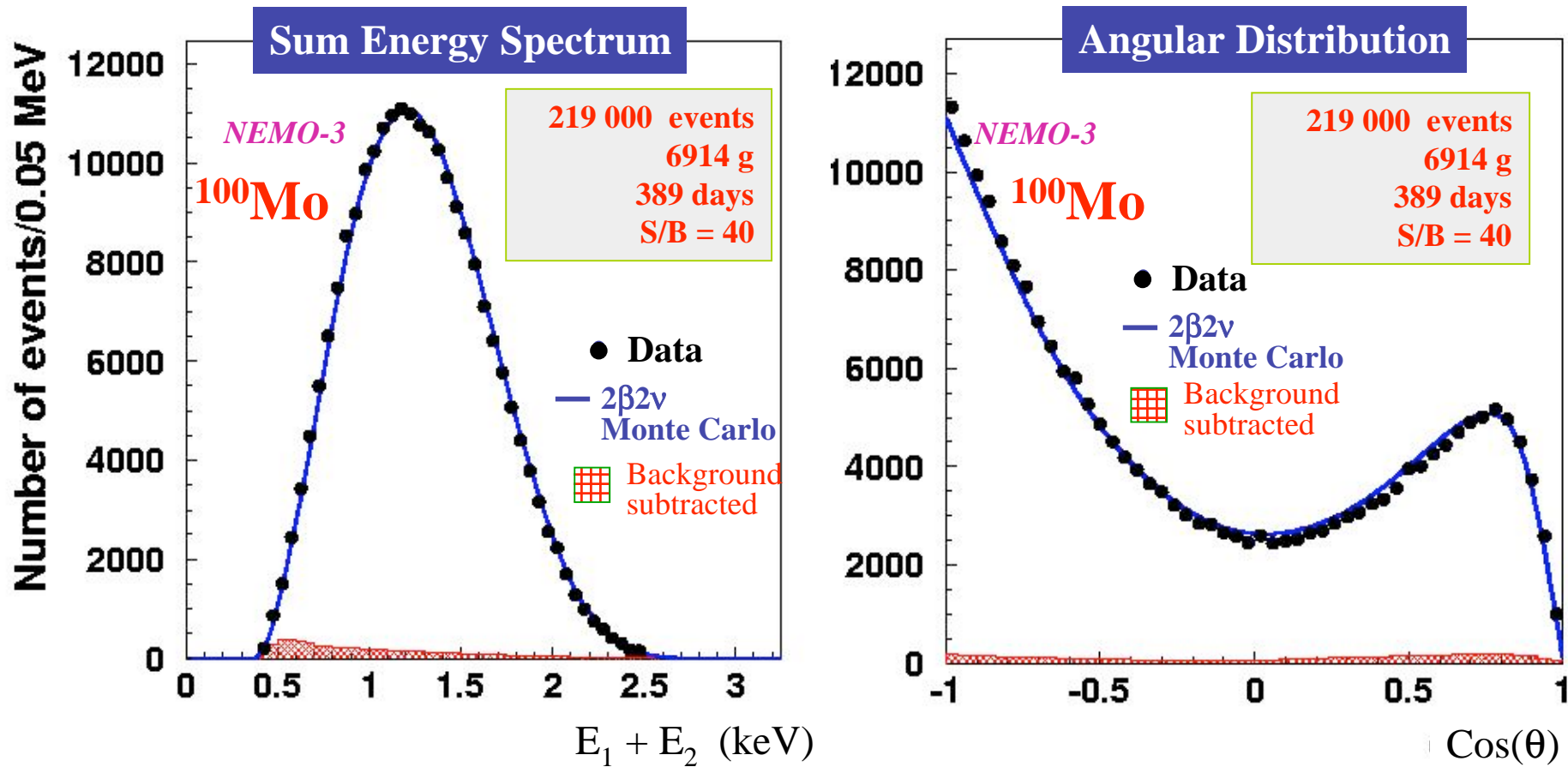


Trigger: at least 1 PMT > 150 keV
 ≥ 3 Geiger hits (2 neighbour layers + 1)

Trigger rate = 7 Hz

^{100}Mo $2\beta 2\nu$ result

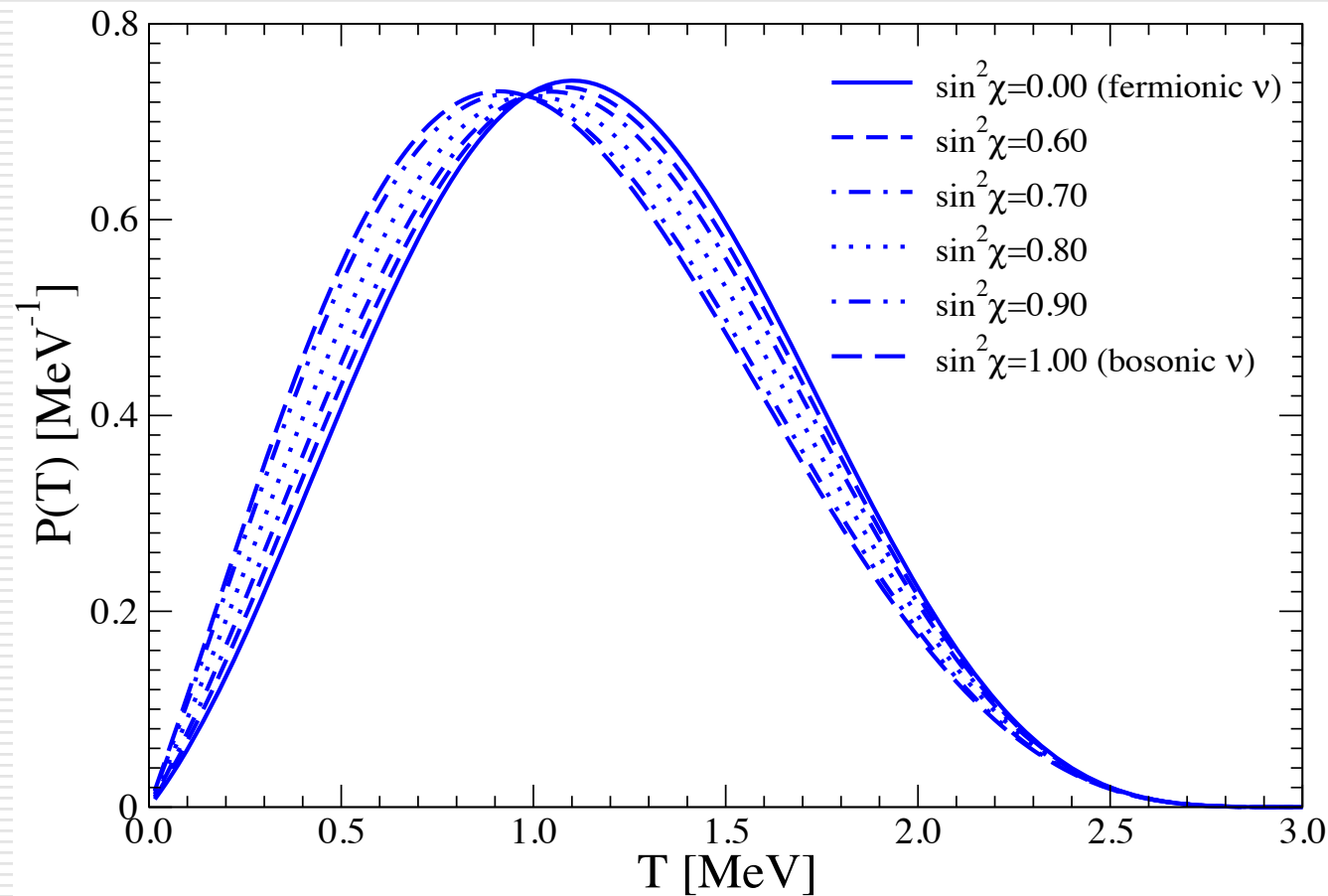
(Phase I: Feb. 2003 – Dec. 2004)



7.37 kg.y

$T_{1/2} = 7.11 \pm 0.02$ (stat) ± 0.54 (syst) $\times 10^{18}$ y

The normalized distribution of the total energy of two electrons



Large admixture of bosonic ν is excluded: $\text{sin}^2 \chi < 0.6$

Conservative limit is $\sin^2\chi < 0.6$
(In fact, best fit gives $\sin^2\chi \approx 0.4-0.5$)

- This is because of the existing disagreement between the data and Monte Carlo simulations (of course, it can be just systematic effect).
 - **NEMO-3** is current experiment. Systematic will be reduced and sensitivity to *bosonic neutrino* will be improved down to $\sin^2\chi \approx 0.2$ (or **bosonic nature of neutrino will be established!**).
-

^{100}Mo . Decay to 2^+ excited state

$$\square T_{1/2}(2^+_1) = 1.7 \cdot 10^{23} \text{ y (fermionic } \nu) \\ = 2.4 \cdot 10^{22} \text{ y (bosonic } \nu)$$

$$r_0(2^+_1) = 7.1$$

Best present limit is $> 1.6 \cdot 10^{21} \text{ y}$

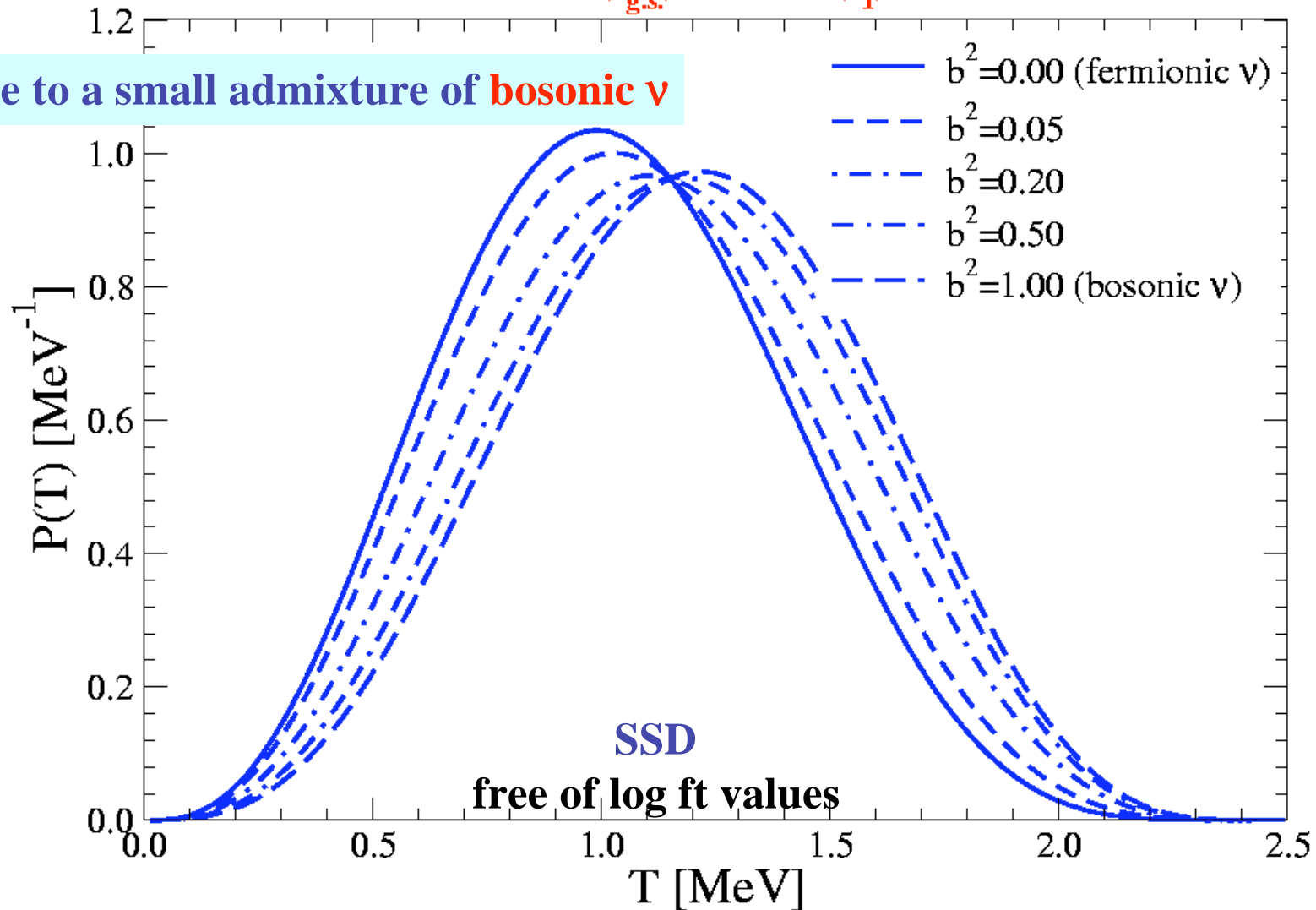
Final sensitivity of NEMO-3 is $\sim 10^{22} \text{ y}$

Sensitivity of SuperNEMO is $\sim 10^{23} \text{ y}$

The normalized distributions of the total energy of two electrons



Sensitive to a small admixture of **bosonic ν**



HSD nuclei even more sensitive to admixture of bosonic ν
but calculation of M.E. needed

CONCLUSION

- A search for anomalous carbon atoms $^{12}\hat{\text{C}}$, atoms with three *K*-shell electrons, is done. A limit on the existence of such atoms was determined: $^{12}\hat{\text{C}}/^{12}\text{C} \leq 2.5 \cdot 10^{-12}$.
 - It corresponds to a limit on the lifetime with respect to violation of **PEP** by electrons in carbon atom of $\tau \geq 2 \cdot 10^{21}$ y.
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CONCLUSION (continuation)

- The **PEP** was tested with the **NEMO-2** detector:

$$\begin{aligned} T_{1/2} &> 4.2 \cdot 10^{24} \text{ y } (^{12}\text{C} \rightarrow ^{12}\hat{\text{C}} + \gamma) \\ &> 3.1 \cdot 10^{24} \text{ y } (^{12}\text{C} \rightarrow ^{12}\tilde{\text{N}} + e^- + \nu) \\ &> 2.6 \cdot 10^{24} \text{ y } (^{12}\text{C} \rightarrow ^{12}\text{B}' + e^+ + \nu) \end{aligned}$$

- Using **NEMO-3** and **SuperNEMO** detectors sensitivity could be increased to $\sim 10^{27}\text{-}10^{28} \text{ y}$
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CONCLUSION (continuation)

- First time **PEP** was checked for neutrino
 - **Pure bosonic** neutrinos are excluded by the present **2β** decay data
 - The existing **2β** decay data allow to put the conservative upper bound
 $\sin^2\chi < 0.6$
 - Sensitivity could be increased up to \sim **0.05-0.1** for **$\sin^2\chi$** (or we will see the "positive" effect)
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